

THERMAL RADIATION WORKSHOP (2012)

RIKEN BNL Research Center Workshop
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Viscosity and thermal photon production

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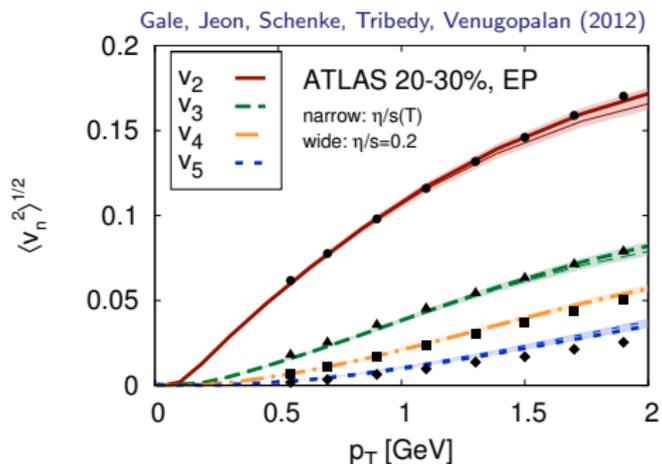
December 7th, 2012

Contents

- ▶ Reminder on shear viscosity and photons
- ▶ Bulk viscosity and hadronic observables
- ▶ Bulk viscosity and electromagnetic observables

Elliptic flow from dissipative hydrodynamics

State of the art fits suggest $\eta/s \sim 0.2$



Still a number of uncertainties (two that will be addressed here)

- ▶ Need for kinetics
- ▶ Bulk viscosity

Elliptic flow from dissipative hydrodynamics

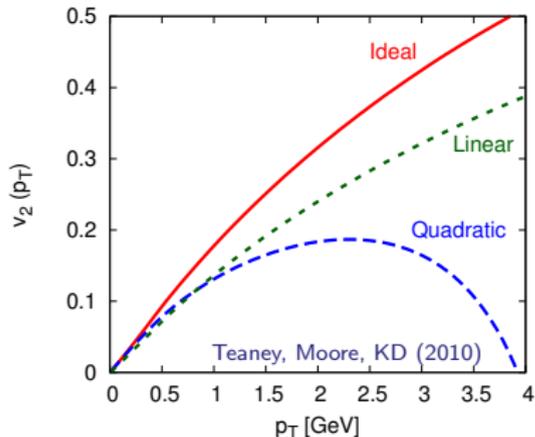
Spectra computed with Cooper-Frye:

$$E_{\mathbf{p}} \frac{dN}{d^3p} = \frac{1}{(2\pi)^3} \int_{\sigma} f(E_{\mathbf{p}}) p^{\mu} d\sigma_{\mu}$$

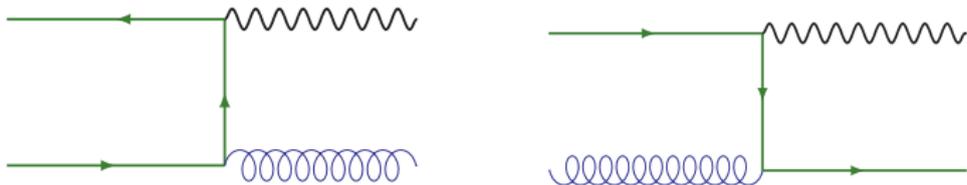
Expand $f = f_o + \delta f$ with constraint:

$$\delta T^{\mu\nu} = \int \frac{d^3\mathbf{p}}{(2\pi)^3 E_{\mathbf{p}}} p^{\mu} p^{\nu} \delta f(E_{\mathbf{p}})$$

Only moments of δf fixed by hydro;
leaving a need for kinetic models.

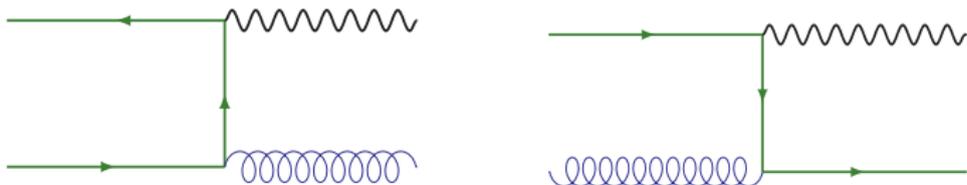


QGP distribution functions



- ▶ Photons are completely out of equilibrium (contrast to early universe)
- ▶ Photon spectra only appears thermal because quarks / gluons creating photons are thermal
- ▶ This is very clear at leading log where $P_{\text{quark}}^{\mu} \approx Q_{\text{photon}}^{\mu}$

QGP distribution functions



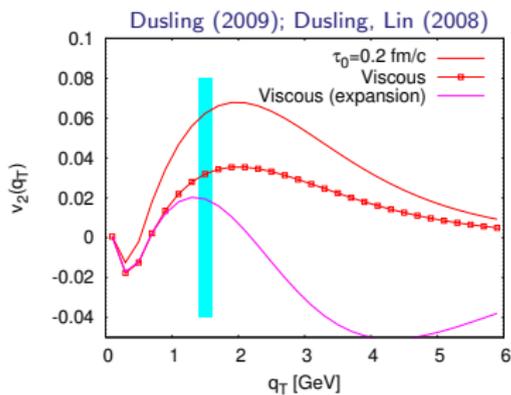
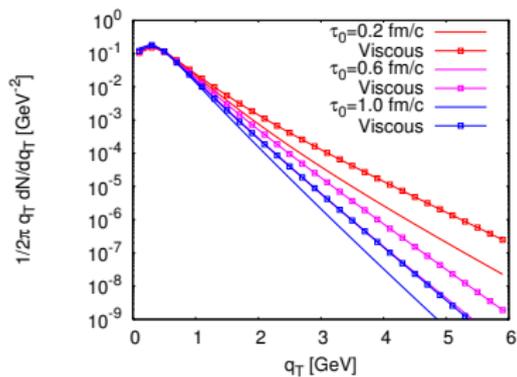
At leading log:

$$E \frac{dN_\gamma}{d^3q} \sim \alpha_{\text{em}} \alpha_S f_{\text{quark}}(Q_\gamma) T^2 \log \left(\# \frac{E_\gamma}{g^2 T} \right)$$

Out of equilibrium:

$$f_{\text{quark}} = f_0 \left(1 - \chi_{\text{quark}}^{\text{shear}}(q) \cdot q^i q^j \partial_{\langle i} u_{j \rangle} - \chi_{\text{quark}}^{\text{bulk}}(q) \cdot \partial_i u^i \right)$$

QGP photons

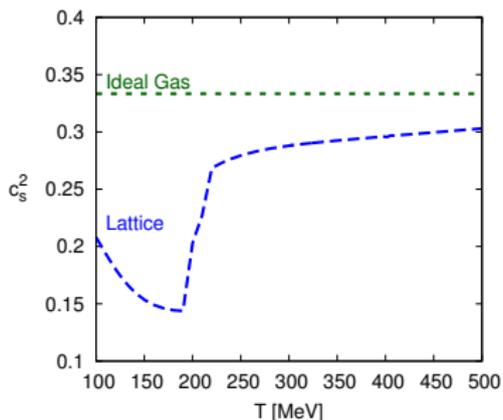


Above calculation shows the kinetic δf correction.

See the more sophisticated work by the McGill group.

Why bulk viscosity?

QCD is clearly not scale invariant and $\zeta \neq 0$.



So, do we need to understand bulk viscosity if we want to extract η/s ?

Yes, but bulk viscosity is interesting in its own right.

Has implications for cosmology (e.g. relic abundances).

Relaxation Time Approximation

Approximate collision operator by single relaxation time:

$$\mathcal{C}[\delta f] \simeq -\frac{\delta f}{\tau_R(E_{\mathbf{p}})}$$

Bulk viscosity goes as 2nd power of conformal breaking

$$\zeta \sim \eta \left(\frac{1}{3} - c_s^2 \right)^2$$

Weinberg (1972)

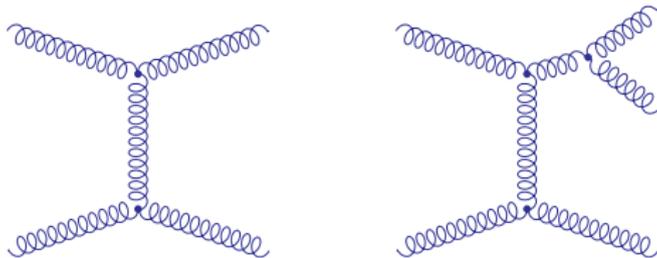
while distribution function goes as 1st power

$$\frac{\delta f}{f_o} \sim p_T^2 \left(\frac{1}{3} - c_s^2 \right) (\partial \cdot u)$$

Bulk viscous correction dominated by δf

QGP distribution functions

QCD: Elastic vs. Inelastic



Arnold, Dogan, Moore (2006)

For elastic $2 \leftrightarrow 2$ recast as Fokker-Planck

$$(\partial \cdot u) \left(\frac{p^2}{3} - c_s^2 E_{\mathbf{p}} \frac{\partial(\beta E_{\mathbf{p}})}{\partial \beta} \right) = \frac{T \mu_A}{n_{\mathbf{p}}} \frac{\partial}{\partial p^i} \left(n_{\mathbf{p}} \frac{\partial}{\partial p^i} \left[\frac{\delta f_{\mathbf{p}}}{n_{\mathbf{p}}} \right] \right) + \dots$$

Drag coefficient:

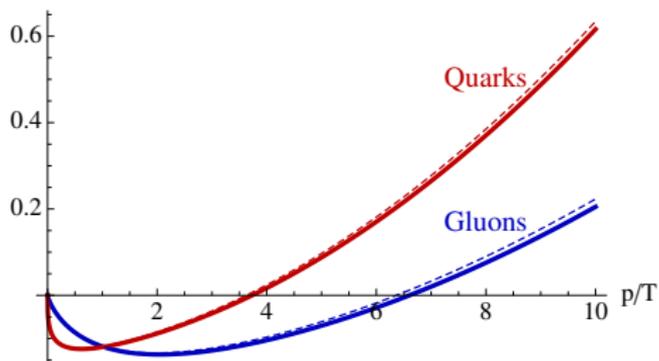
$$\mu_A = \frac{g^2 C_A m_D^2}{8\pi} \ln \left(\frac{T}{m_D} \right)$$

QGP distribution functions

Result:

$$\delta f_{\text{bulk}} \sim \left(\frac{1}{3} - c_s^2 \right) \delta f_{\text{shear}} \quad \zeta \sim 50 \left(\frac{1}{3} - c_s^2 \right)^2 \eta$$

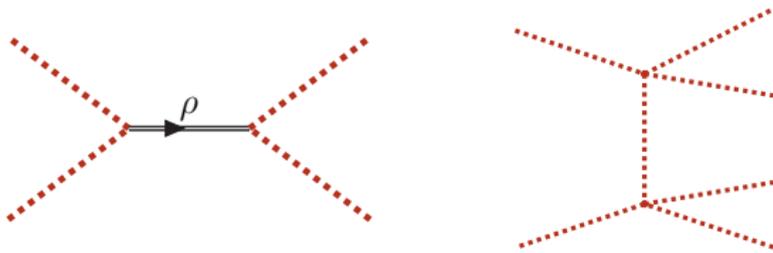
$\chi(p)$



where $\delta f = -f_o (\partial \cdot u) \chi(p)$

Pion Gas:

Pion Gas: Elastic vs. Inelastic



Bulk viscosity governed by chemical non-equilibrium

δf takes form of zero mode which dominates C^{-1}

$$\begin{aligned}\delta f &= -f_o (\chi_0 - \chi_1 E_{\mathbf{p}}) (\partial \cdot u) \\ &= f_o \left(\frac{\delta \mu}{T} + \frac{\delta T}{T^2} E_{\mathbf{p}} \right)\end{aligned}$$

Pion Gas:

Bulk viscosity governed by chemical non-equilibrium

$$\delta f = -f_o (\partial \cdot u) (\chi_0 - \chi_1 E_{\mathbf{p}})$$

Chemical equilibration rate determines χ_0 , energy conservation fixed χ_1

$$\chi_0 = \frac{\beta \mathcal{F}}{4\Gamma_{2\pi \rightarrow 4\pi}}, \quad \zeta = \frac{\mathcal{F}^2}{4\Gamma_{2\pi \rightarrow 4\pi}}$$

where \mathcal{F} characterizes conformal breaking

$$\mathcal{F} \equiv \int \frac{d^3 \mathbf{p}}{(2\pi)^3 E_{\mathbf{p}}} \left(\frac{p^2}{3} - c_s^2 E_{\mathbf{p}}^2 \right) f_o$$

Hadron Resonance Gas: A model

Assume slowest process is chemical relaxation:

$$\delta f^a = -f_o (\partial \cdot u) (\chi_0^a - \chi_1 E_{\mathbf{p}})$$

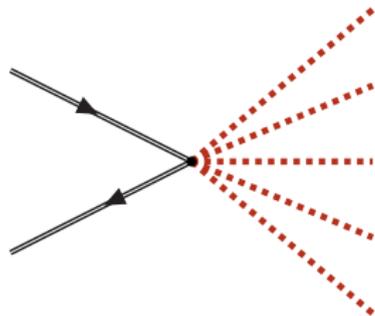
where $a = \pi, K, \rho, K^*, p, n, \Delta, \dots$

Slowest rate determines ζ , other rates fix the relative δf^a :

$$\chi_0^a \simeq \chi_0^\pi \times \begin{cases} 2 & \text{Mesons} \\ 2.5 & \text{Baryons} \end{cases}$$

Motivated by $\mu_\rho = 2\mu_\pi$ and $2\mu_N = 5\mu_\pi$:

Goity (1993); Pratt, Haglin (1999)



Hadron Resonance Gas: A model

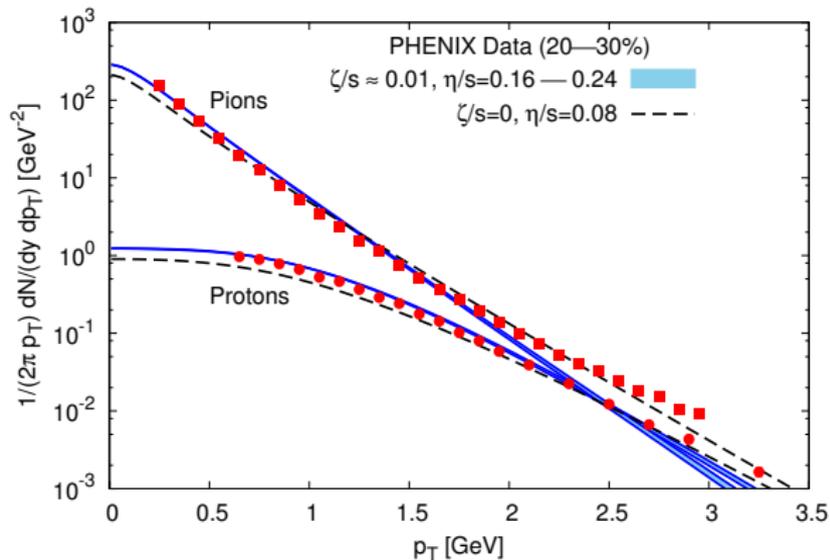
Hadron Gas: ζ determined by chemical non-equilibration

Bottom line for δf :

- ▶ $\delta\mu \sim \frac{\zeta}{s} (\partial \cdot u)$
- ▶ temperature shift $\delta T \approx 0.25\delta\mu$ to conserve energy

A new (dynamical) way to look at fugacity factors

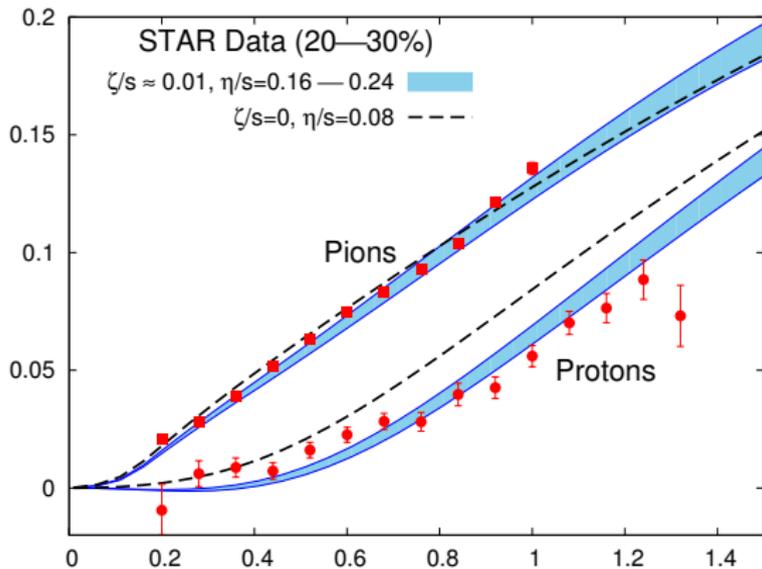
Pion / Proton p_T spectra



Data: PHENIX nucl-ex/0307022.

LHC: Bozek & Wyskiel arxiv:1203.6513

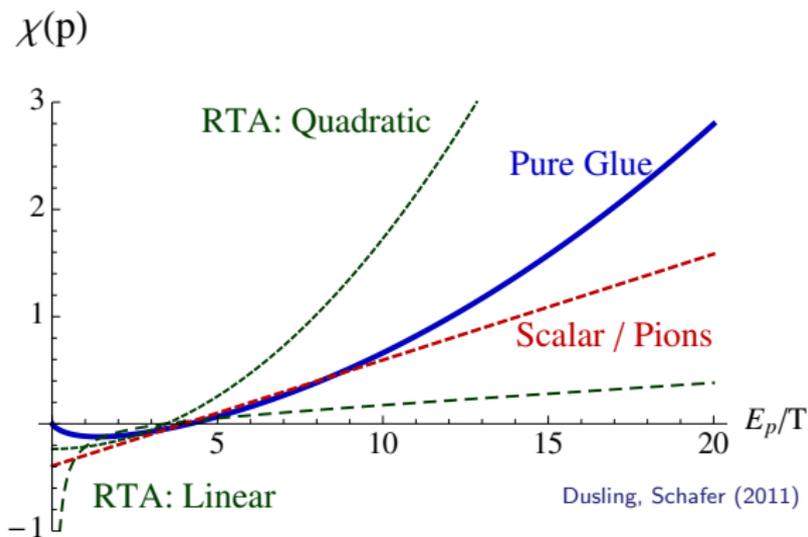
Pion / Proton differential $v_2(p_T)$ spectra



Data: STAR, nucl-ex/0409033.

Bulk viscosity and EM probes

As for shear viscosity, bulk δf modifies photon rates

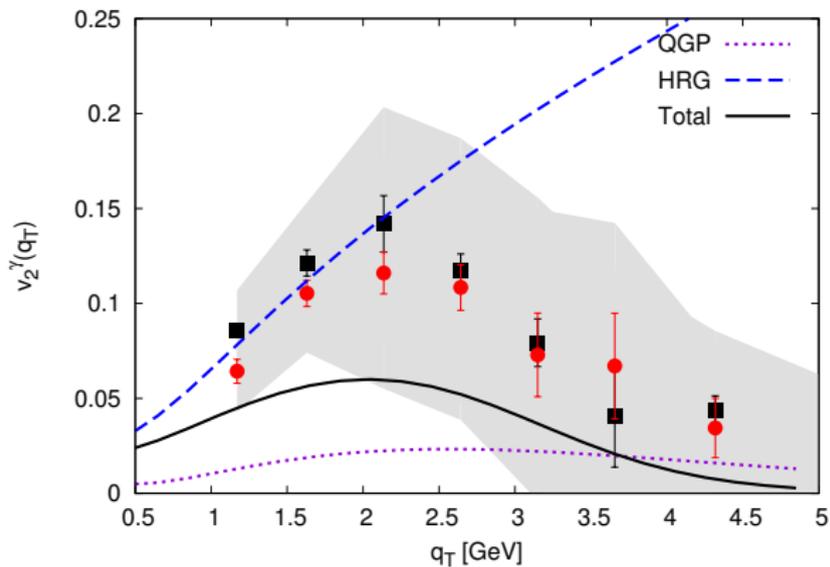


Bulk viscosity and EM probes

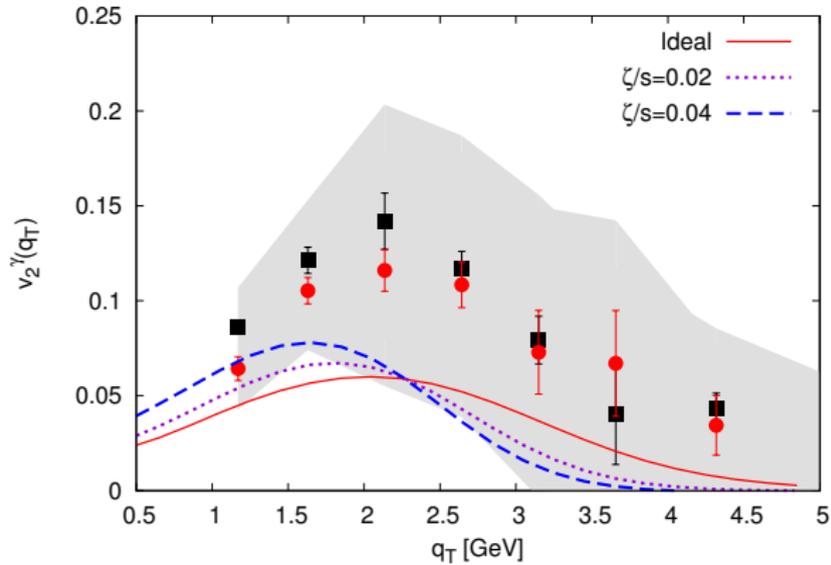
First shot:

- ▶ Using AMY rates with $f \rightarrow f + \delta f$ and pQCD δf
- ▶ Using χ RF (Steele, Yamagishi, Zahed) with
 - ▶ z_π^3 enhancement with $\mu_\pi \sim \zeta/s (\partial \cdot u)$
 - ▶ $T \rightarrow T - \delta T$
- ▶ Results preliminary: Need to
 - ▶ include shear viscosity
 - ▶ fine-tune initial condition
 - ▶ worry about pQCD at high q_T
 - ▶ ...

Ideal photon $v_2(q_T)$



Viscous photon $v_2(q_T)$



Only including bulk viscosity – no fine-tuning.

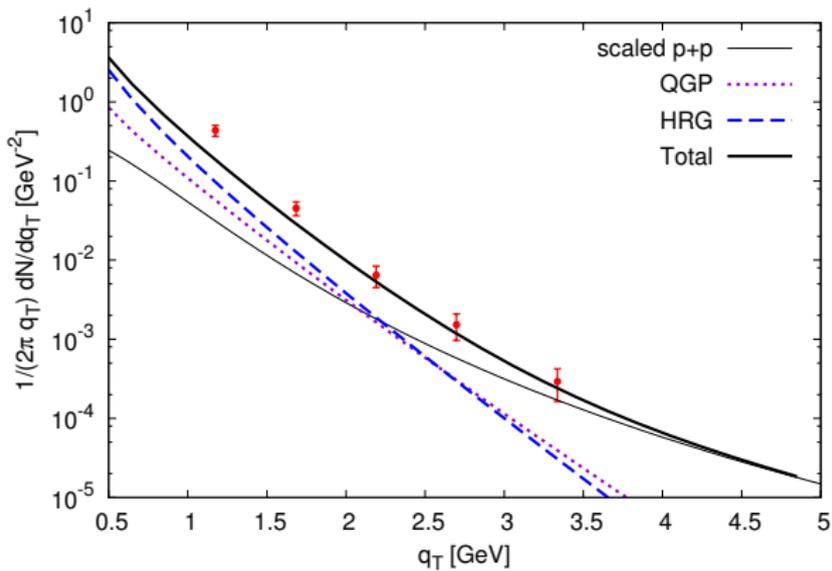
Summary

Bulk viscosity is not zero:

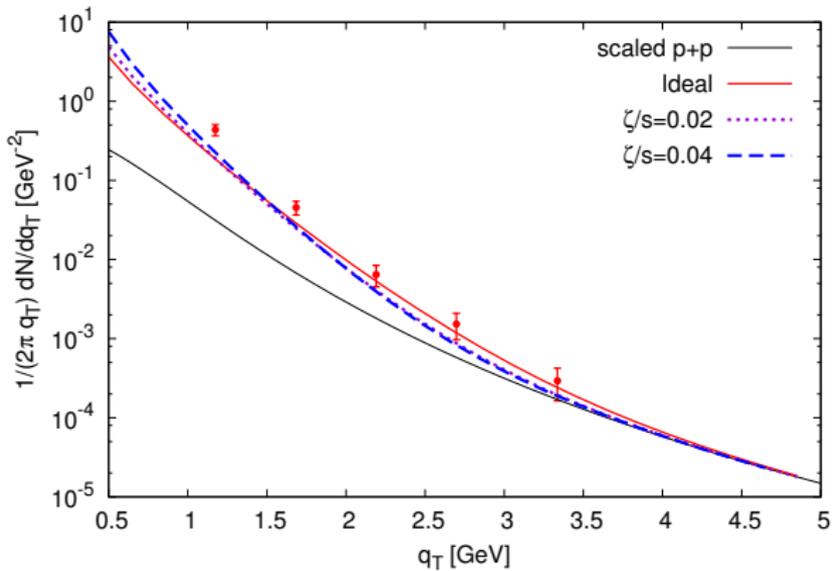
- ▶ fine structure of spectra improves
- ▶ may help with photon v_2
- ▶ dynamical mechanism for fugacity factors

Backup Slides

Ideal photon q_T spectra



Viscous photon q_T spectra



Only including bulk viscosity – no fine-tuning.

Thermal photon production

At leading order in α_{em} but all orders in α_{s}

$$\frac{dN}{d^4Q} = \frac{\alpha_{\text{em}}^2}{6\pi^3} \frac{1}{Q^4} (Q^\mu Q^\nu - Q^2 g^{\mu\nu}) W_{\mu\nu}(Q)$$

where

$$W_{\mu\nu}(Q) \equiv \int d^4x e^{-iQ \cdot X} \langle J_\mu^{\text{em}}(X) J_\nu^{\dagger, \text{em}}(0) \rangle_\beta$$

Evaluate $W_{\mu\nu}$ in two different ways

- ▶ Vacuum spectral functions
- ▶ Kinetic Theory

Thermal photon production

Kinetic theory (all processes $I \rightarrow F + l^+l^-$):

$$W_{\mu\nu}(Q) = \sum_F \sum_I \int d^4x e^{-iQ \cdot X} \langle I | J_\mu^{\text{em}}(X) | F \rangle \langle F | J_\nu^{\dagger, \text{em}}(0) | F \rangle \frac{e^{-\beta E_I}}{\mathcal{Z}}$$

using $E_I = E_F + Q_0$ and $\sum_I |I\rangle \langle I| = 1$

$$W_{\mu\nu}(Q) = e^{-\beta Q_0} \sum_F \int d^4x e^{+iQ \cdot X} \langle F | J_\mu^{\dagger, \text{em}}(X) J_\nu^{\text{em}}(0) | F \rangle \frac{e^{-\beta E_F}}{\mathcal{Z}}$$