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#### Gene Van Buren Brookhaven National Lab

**STAR Experiment** 

#### Co-authors: L. Didenko, J. Lauret (BNL), J.H. Thomas (Lawrence Berkeley National Lab)

# Adaptive on-the-fly calibration of TPC distortions

### **STAR TPC**



### **TPC Distortion Corrections**

- Each TPC distortion correction requires some "measure" of the problem:
  - Field maps, surveys, reconstructed track observables
  - Observables are most easily determined from some set of "ideal" tracks (e.g. perfectly straight) which may require large statistics (many reconstructed events)
- Most distortions have static causes
- Conflict: what to do when a highly dynamic (volatile) distortion needs large stats? .....

### SpaceCharge: model of charge

HIJET model of "event shape" for 200 GeV AuAu collisions matches radial distribution of zerobias data well for much of the runs.



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March 1, 2004 data

### **Distortion equations**

(see Blum & Rolandi)

Solve:

$$m \frac{d\overline{u}}{dt} = e \overline{E} + e \left[\overline{u} \times \overline{B}\right] - K \overline{u}$$

substituting:

Langevin Equation with "Friction"

$$\tau = \frac{m}{K}$$
,  $\omega = \frac{e}{m} |\overline{B}|$ ,  $\mu = \frac{e}{m}\tau$ , and  $\hat{E} = \frac{E}{|\overline{E}|}$ 

subject to the steady state condition

$$\frac{d\overline{u}}{dt} = 0 \qquad \text{yields}$$

$$\overline{u} = \frac{\mu |\overline{E}|}{(1+\omega^2\tau^2)} \left( \hat{E} + \omega\tau \left[ \hat{E} \times \hat{B} \right] + \omega^2\tau^2 \left( \hat{E} \cdot \hat{B} \right) \hat{B} \right)$$

If you have a well defined model, and good data, then the distortion can be removed with great precision







"I think you should be more explicit here in step two."

If you have a well defined model, and good data, then the distortion can be removed with great precision

on"

 $\overline{E}$ 

 $|\overline{E}|$ 

 $(\hat{B})\hat{B}$ 

### **SpaceCharge Field Effects**

- Using our "event shape" model
  - Relaxation done on 5cm x 5cm 2D
     (r-z) grid (assume Φ symmetry)
  - Treat as a perturbation on top of standard TPC E field
  - Distortions are integral of E field in z (drift direction)
  - Not very sensitive to radial component of distortion because tracks are radial-like
  - Lorentz Force Eqn:  $\vec{F} \propto q \cdot (\vec{E} \times \vec{B})$



#### **SpaceCharge effect on sDCA** All tracks go the same direction VEXB (pos. or neg.) Ve Ve**∧** B out of Track charge plane **V**EXB independence Field dependence DCA Positive physical sDCA sDCA = signed distance of closest approach

### **SpaceCharge effect on sDCA**



### **TPC GridLeak distortion**



 Dependence on field, track charge, location, luminosity consistent with ion leakage at gating grid gap

### **TPC GridLeak distortion**



ion leakage at gating grid gap

### **GridLeak Field Effects**

#### Modeled sheets of charge

- Relaxation done on custom 3D grid
   (plots assume Φ symmetry, but leak is 12-fold symmetry from grid shape)
- E-field and distortion discontinuity at grid gap

#### GridLeak scales as SpaceCharge!







### **Applied GridLeak Correction**

Not perfect, but as good as Residual [cm] 0.04 design spec! 0.02 0 **Distortions scale** -0.02 significantly reduced! -0.04 -0.06 -0.08+ 200 180 160 50 140 R [cm] 120 100 80 **60** 0 Residual [cm] 0.04 0.02 0 -0.02 After -0.04 200 150 [cm] -0.06 Before -0.08 100 200 180 160 50 140 120 R [cm] 100 80 60 0

200

150 [cm]

100

### First steps to corrections

 Observables (sDCA) can tell you the distortion quantity (ions in the TPC due to SpaceCharge buildup + GridLeakage)

#### Easy with "ideal" tracks

- Little or no dependencies on reconstruction itself
- Observable maps easily to distortion quantity
  - SDCA = C \* f(Z) \* (SpaceCharge + GridLeak)

Generally need many events for stats

Could be many <u>runs</u> for pp collisions!

### First steps to corrections

Observables (s distortion quar SpaceCharge k
Easy with "idea
Little or no dependent

Observable map



SDCA = C \* f(Z) \* (SpaceCharge + GridLeak)

Generally need many events for stats

Could be many <u>runs</u> for pp collisions!

### **Ionization: Scalers**

- Ionization is linear with scaler measures<sup>0.014</sup>
   of luminosity
- Points out
   problem runs
- Some smearing from 30 second average



STAR records scaler rates on Zero Degree Calorimeters (ZDCs) and Beam-Beam Counters (BBCs)

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 Collision rates look smooth in ZDC rates...

...but
 background
 rates show
 something
 going on...



Fill 4529, February 13, 2004

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Fill 4547, February 16, 2004

 Collision rates look smooth in ZDC rates...

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Fill 4547, February 16, 2004



### **Event-by-Event: use history**

Use several recent events
 with age weighting

Throw and refit 10
 tracks with simple (quick) 1
 model to allow for larger 1
 selection



### **E-by-E Successes**



Differences show method uncertainty

### **E-by-E Successes**



### **E-by-E Successes**



### **E-by-E Issues**



### **E-by-E Issues**



### Performance Measures: sDCA

- Can't beat low luminosity, but holding steady at high luminosity:
  - Spread from
     5-9kHz appears
     roughly uniform
  - No indication we can't go higher!

2004 AuAu at 200 GeV, all B fields



### **Performance Measures:** $\pi^{-}/\pi^{+}$

 TPC-measure of the ratio
 essentially flat
 all the way to
 pT=12 GeV/c !

 Central triggers (taken at high luminosity) just about as good!



## Summary & Outlook

Developed a technique to determine and apply TPC distortion corrections on an event-by-event basis No obvious luminosity limitations Possible improvements: Generation Higher frequency scalers (instead of Prepass) Address backgrounds (shielding, 3D correction) Searce Apply correction to same event (must measure distortion during tracking) So Fixed reference (silicon tracking)