

# EPD SIMULATION: STATUS UPDATE

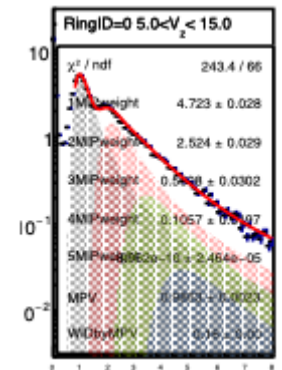
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EPD Meeting Jan. 14 2021

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# Measuring $dN/d\eta$ with the EPD

- EPD measures signal (ADC)  $\rightarrow$  Convolution of several Landau distributions
- With "multiple Landau" fits, one can extract  $dN/dn_{\text{MIP}}$  for each ring
  - See details e.g.:  
<https://drupal.star.bnl.gov/STAR/blog/lisa/extracting-dndeta-forward-region-unfolding>



- Each event has a given hits in a given ring:  $N(i_{\text{Ring}})$
- Originates from an underlying  $dN/d\eta$ ,  $N(i_{\text{Ring}})$  can be calculated as

$$N(i_{\text{Ring}}) = \int R(\eta, i_{\text{Ring}}) \frac{dN}{d\eta} d\eta$$

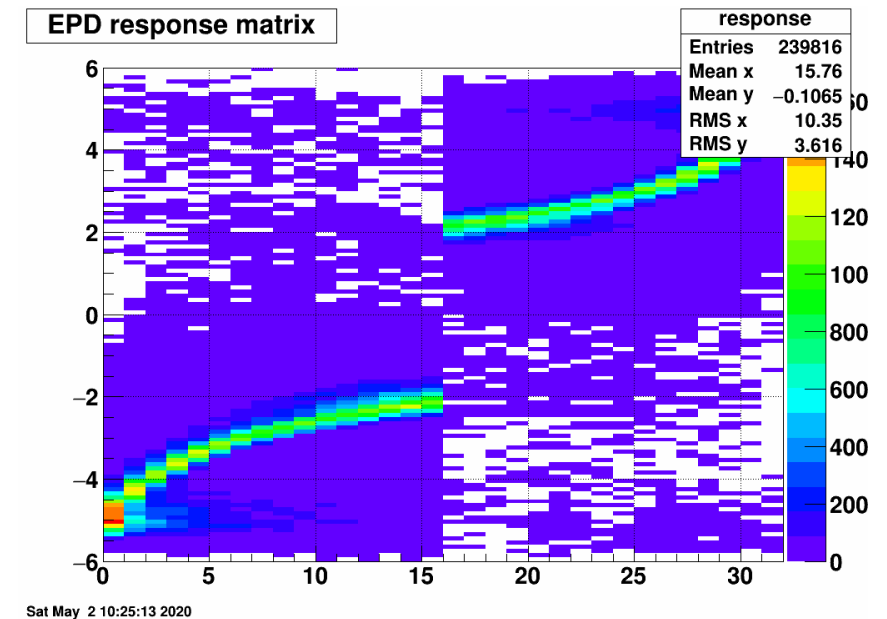
- Here  $R$  is the **response matrix**: how many hits in given ring from particles with  $\eta$
- How to invert this?
- Calculate  $R$  via simulations, determine  $dN/d\eta$  via **unfolding**

# Calculating the response matrix via simulation

- Use iterative unfolding, based on G. D'Agostini, Nucl. Instr. Meth. A362 (1995) 487
- Implemented in RooUnfold, response matrix to be calculated as:

```
for(PrimaryTracks)
{
  if(no EPDhits from that Primary Track)
  {
    R->Miss(TrackEta); //This track "missed" the EPD
  }
  else
  {
    for(EPD hits of that Primary Track)
    {
      R->Fill(EPDeta,Tracketa);
    }
  }
}
```

- Simulation part: need
  - list of primary tracks
  - EPD hits and the primary track that caused them
- All possible in HIJING+GEANT simulator, using StarHijing (1.383 ) and StEpdFastSimMaker
  - Further utilities used: StarGeneratorUtil/Event/Base, St\_geant\_Maker, etc.



# Technical details (how-to) of the simulation

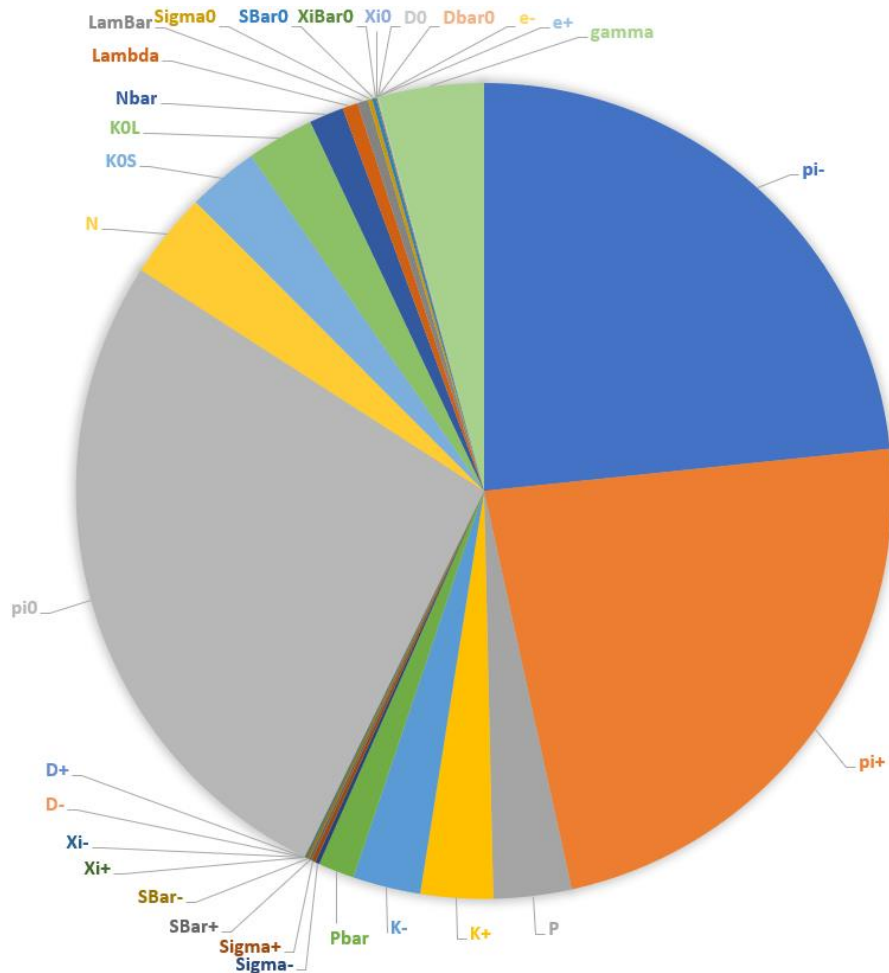
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1. Copy components from `/star/u/mcsanad/newsim/hijingPlusGeantSim/submit/starsim.C`, `runEpdFastSim.C`, `runHijing.xml`, `makeMuDST.xml` (plus `StRoot` directory checked out)
2. Edit xml files to reflect on user directory and username, create `log`, `err`, `out` and `fzdroot` directories (given in xml files)
3. Edit settings in `runHijing.xml`:
  - SL version (SL19e seemed to work)
  - nProcesses (500 currently) and NEVENTS (10 currently)
  - geometry tag (y2018a seemed to work)
  - a random seed for the first job (currently 26544321)(`makeMuDST.xml` needs probably the same SL version)
4. Edit settings in `starsim.C`:

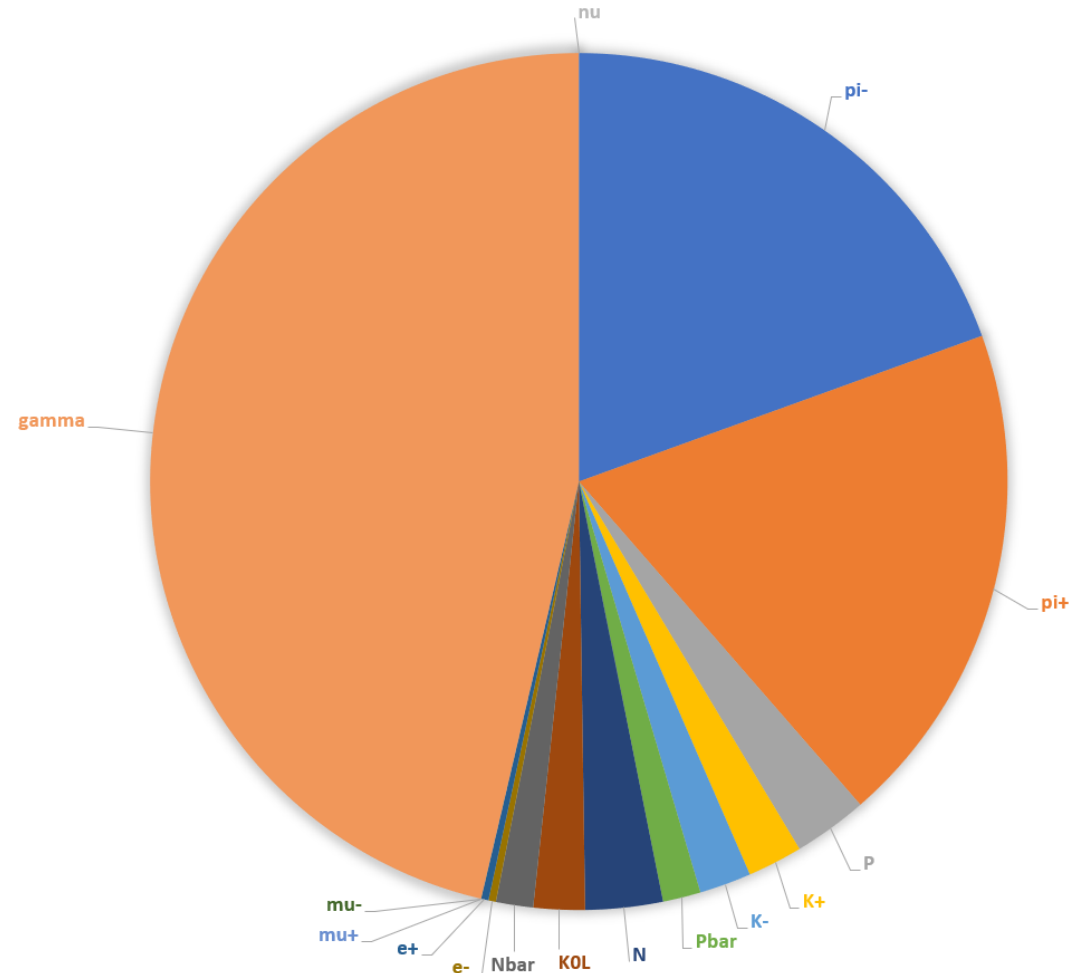
```
hijing->SetFrame("CMS",200.0); //CMS energy per nucleon pair
hijing->SetBlue("Au");
hijing->SetYell("Au");
hijing->SetImpact(0.0, 1.0); // b in [0 fm, 1 fm]
```
5. Submission: `star-submit runHijing.xml` (modify queue if needed)
6. List resulting fzd files in text file, edit `makeMuDST.xml` to use that list in `<input URL="filelist:...">`, then `star-submit makeMuDST.xml`
7. Example analysis codes are in `/star/u/mcsanad/newsim/`  
see for example: `root.exe -b -q lMuDST.C ZvtxBinnedResponse.C+`

# What particles are there in the simulation?

Before decays

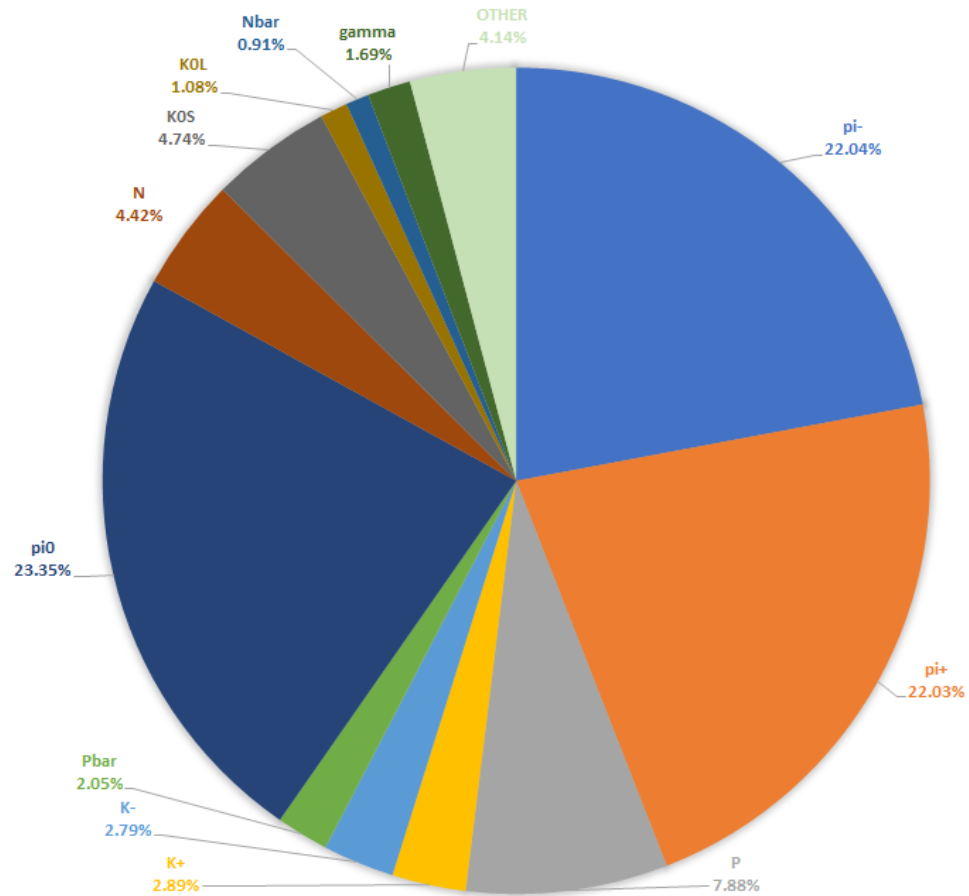


After decays

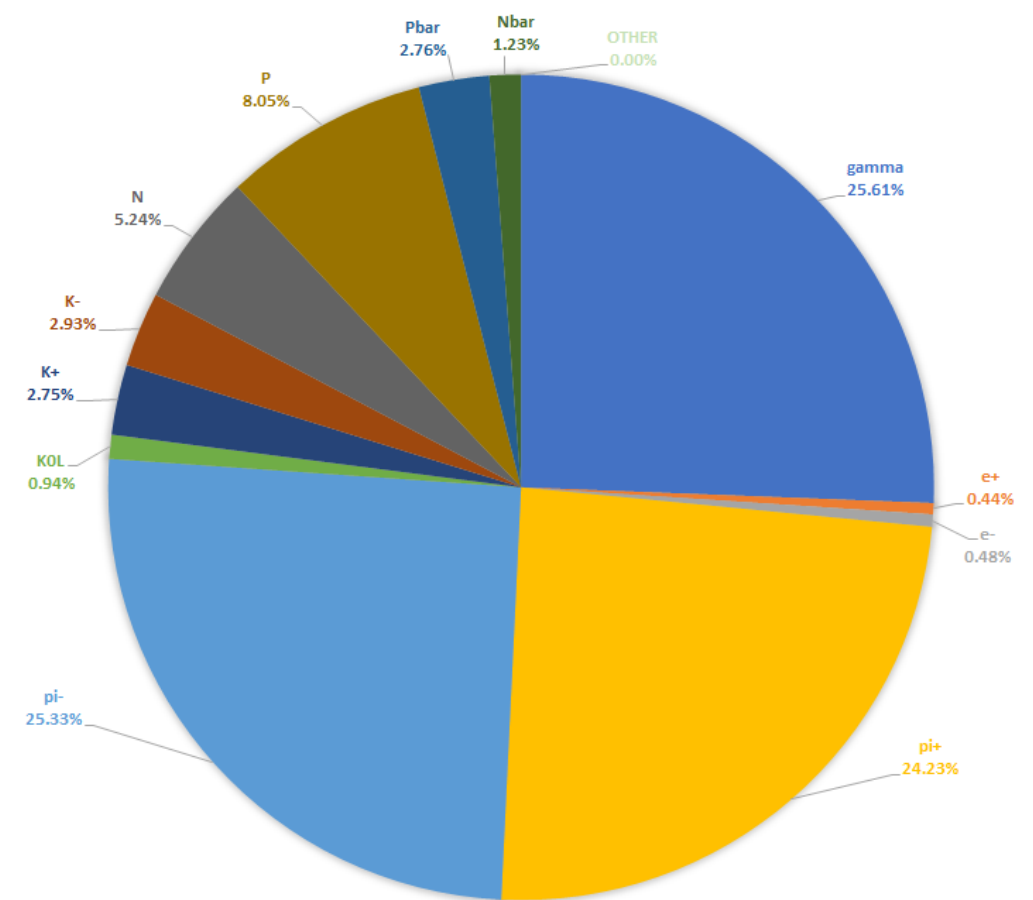


# "Original" particles causing hits in the EPD

Before decays

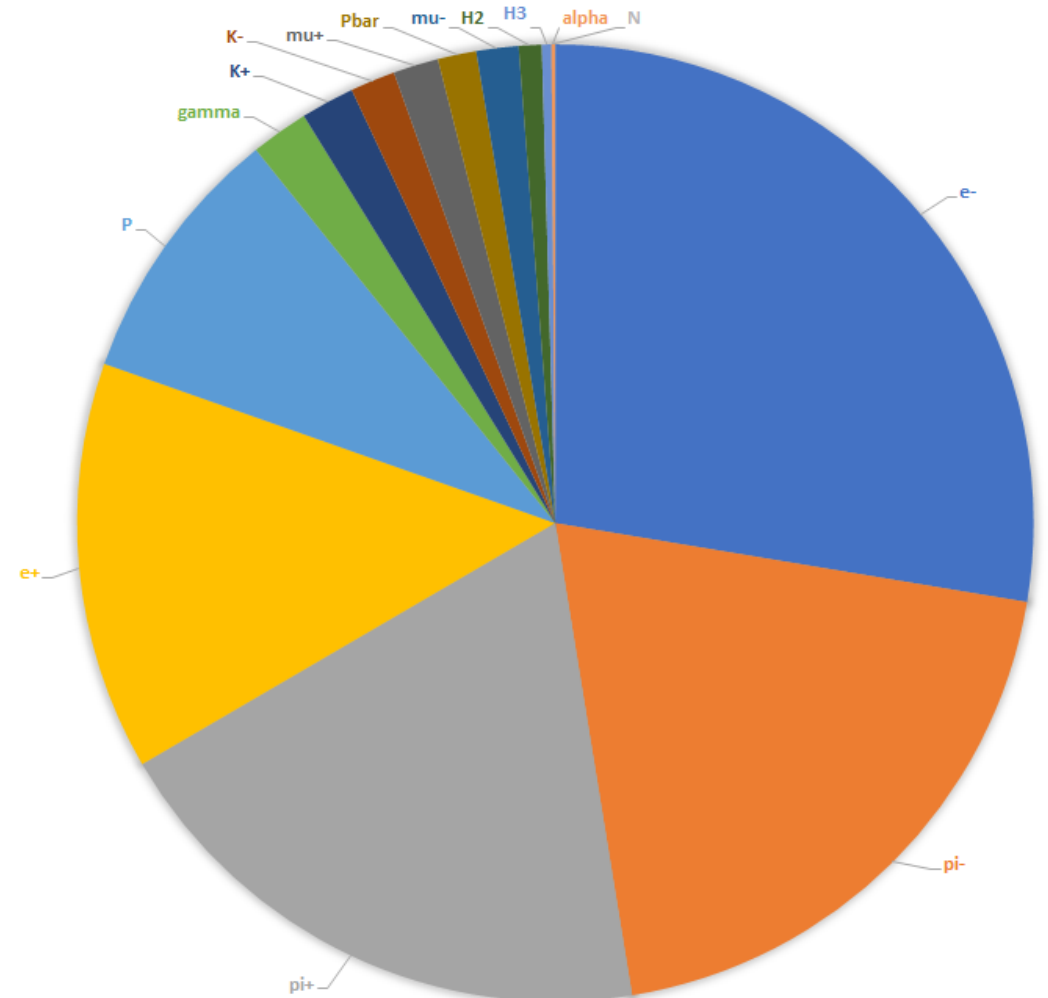
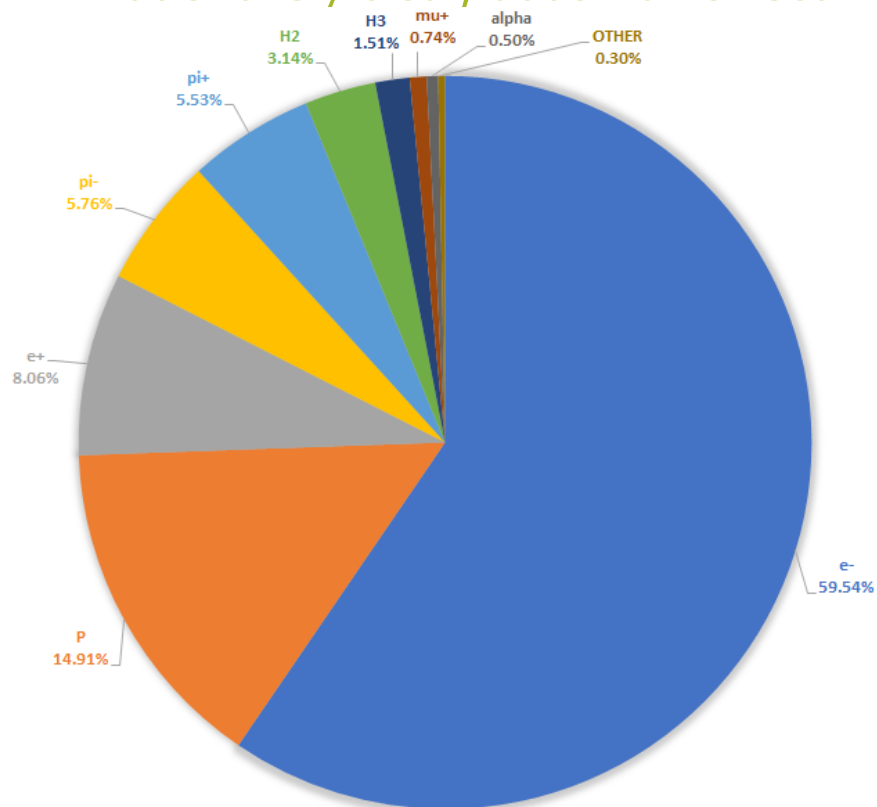


After decays



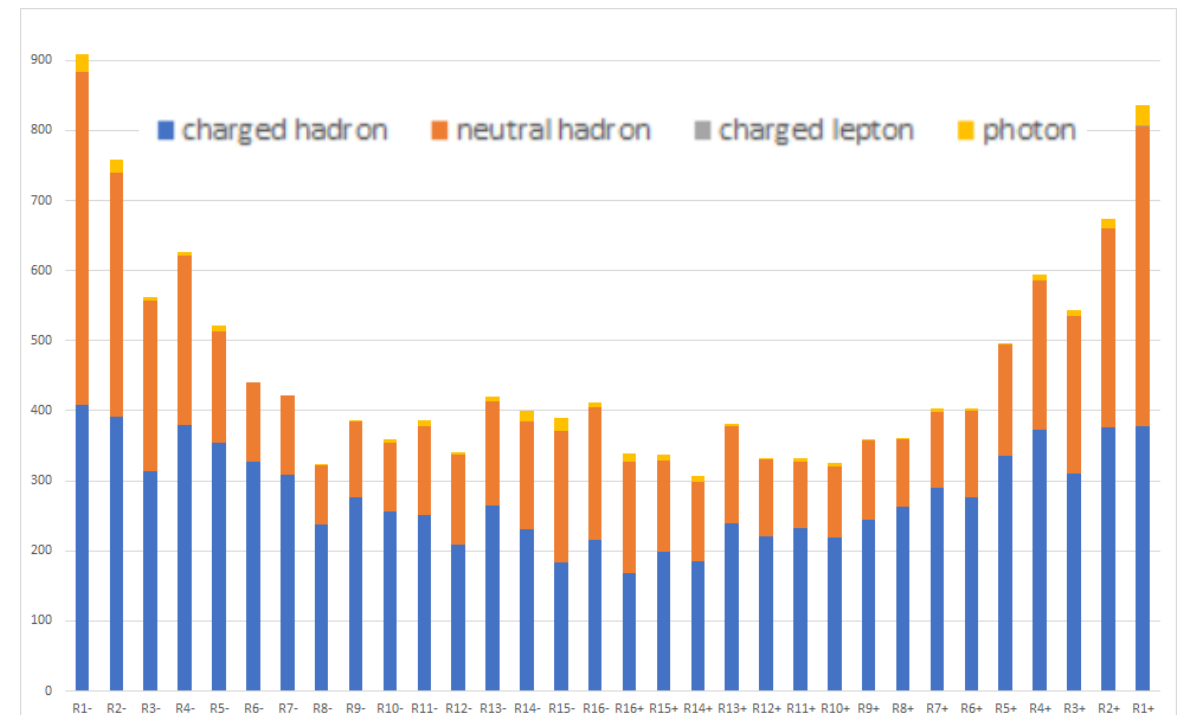
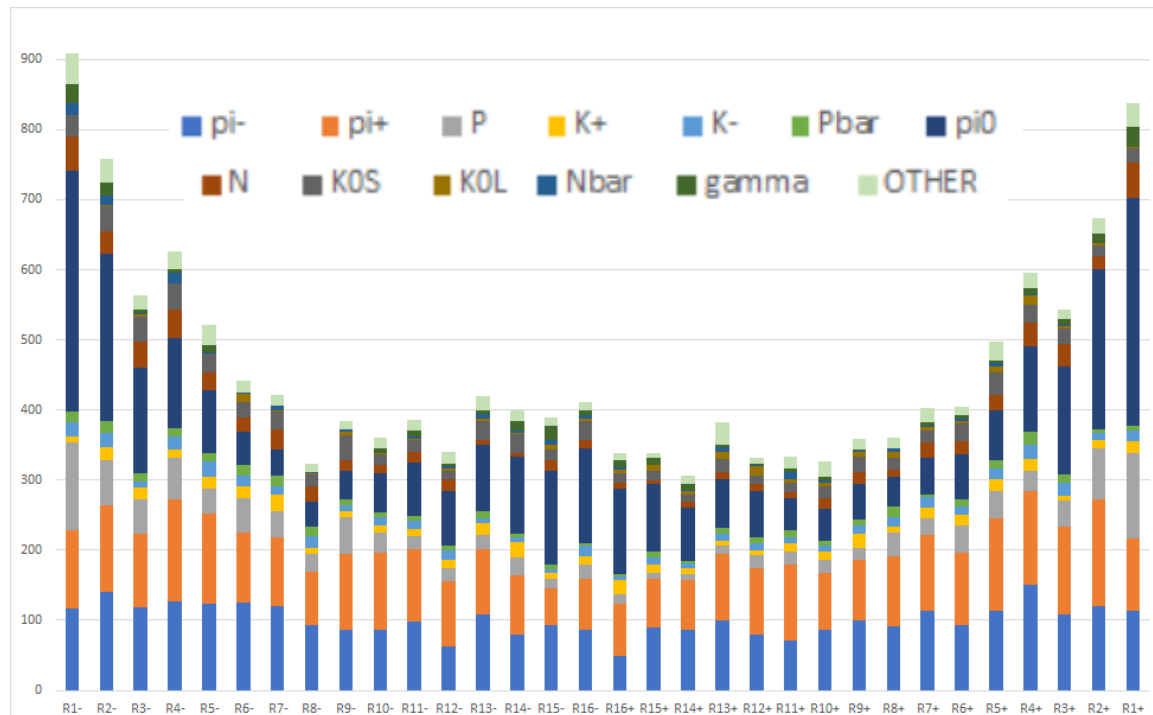
# Particles directly hitting the EPD

- Mostly secondary (i.e. GEANT)
- Some are born inside the EPD
  - Not entirely clear, but small effect



# EPD hit causing primaries, ring-by-ring

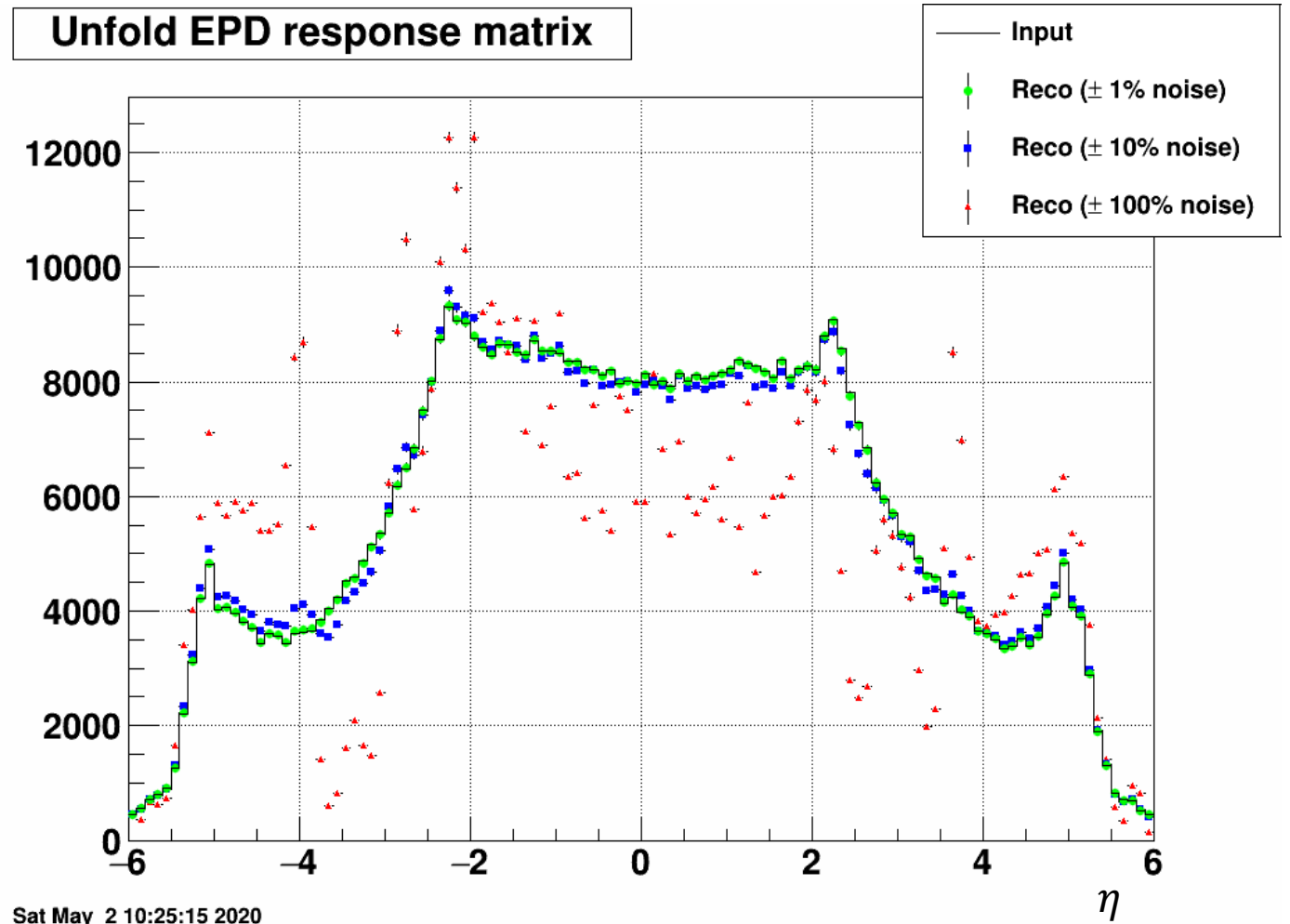
- Fraction of charged hadrons <50% for inner&outer rings, >70% for medium rings
- Rest: neutral hadrons (mostly  $\pi^0$ , neutron,  $K^0$ ), few photons





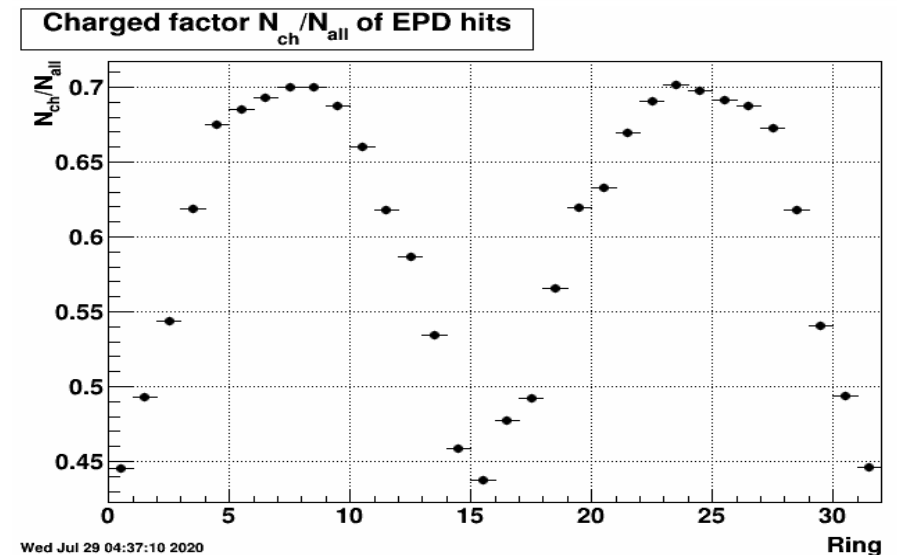
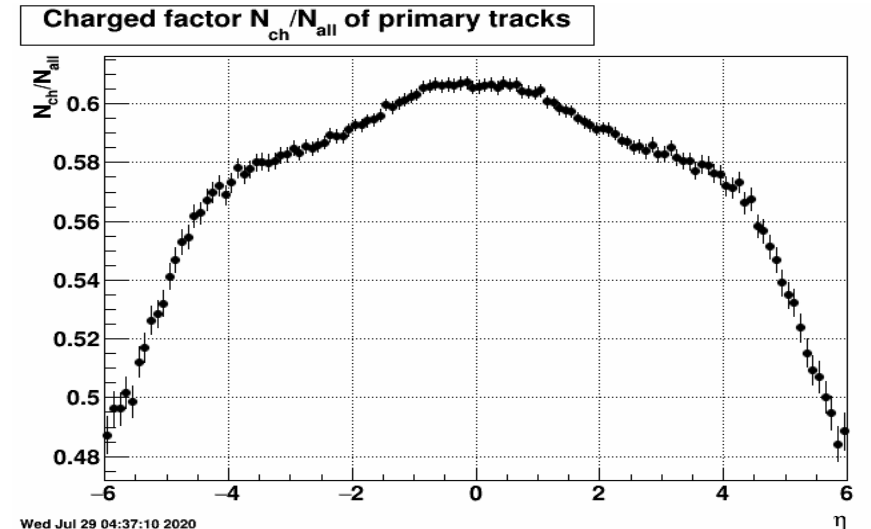
# Does the unfolding work?

- If unfolding on training sample: returns input perfectly
- Adding some noise: imperfect but still good
- Why the peaks near  $\eta = 5$ ?
  - One unfolded track for each individual EPD hit
  - Many tracks cause multiple hits  $\rightarrow$  need to correct for this!
- How can it work near  $\eta = 0$ ?
  - It reconstructs  $dN/d\eta$  of input!
  - Need to investigate systematic uncertainty from input sample



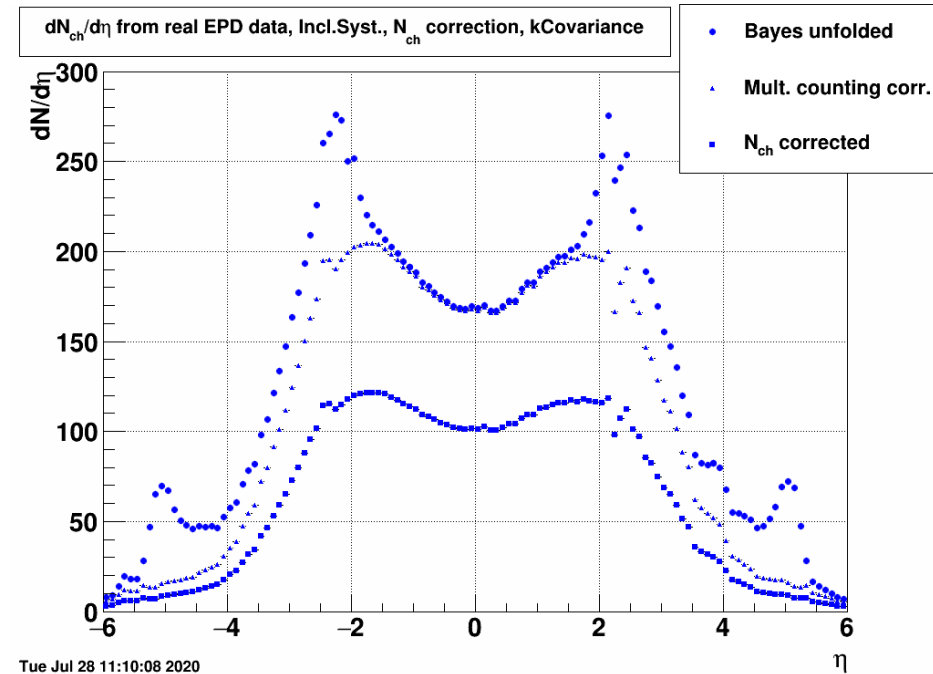
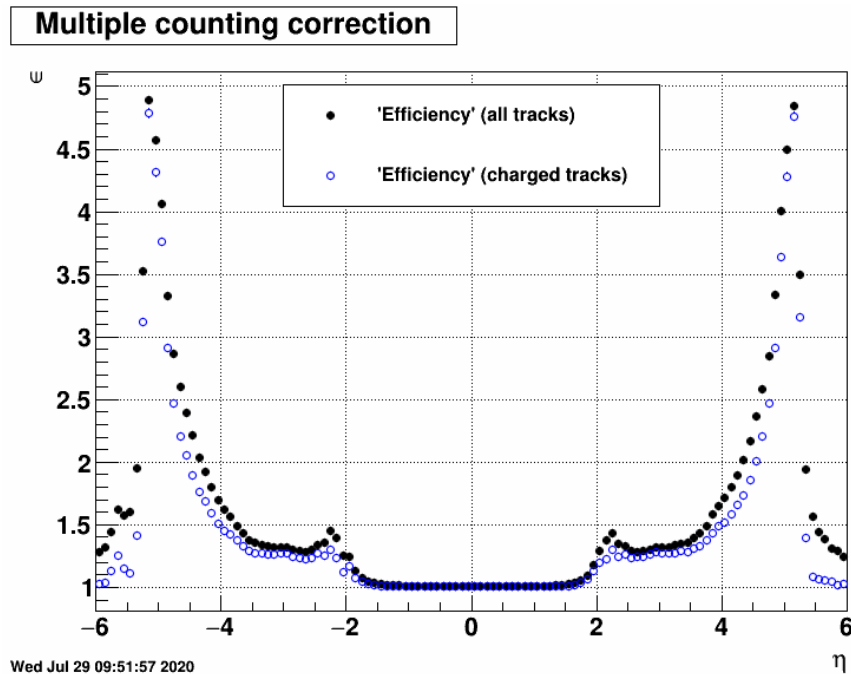
# Measure charged tracks only?

- Known in simulations: charged factor
  - For primary tracks
  - For EPD hits (based on primary cause)
- Tried 3 possible methods:
  1. Unfolding  $dN/d\eta$ ; correcting via  $N_{\text{ch}}(\eta)/N_{\text{tot}}(\eta)$
  2. Correcting via  $N_{\text{ch}}(i_{\text{ring}})/N_{\text{tot}}(i_{\text{ring}})$ ; unfolding "corrected" EPD distribution
  3. Use RooUnfold's "Fakes" (neutrals  $\Leftrightarrow$  "fake" hits)
- First two work well
- Method 3: more dependence on input  $dN/d\eta$
- Difference of methods: incorporate in systematics



# Multiple hits inverse efficiency correction

- Need to correct for multiple counting (many hits from one primary track)
  - Check "inverse efficiency": how many hits on average at given  $\eta$
- Need to unfold charged tracks only

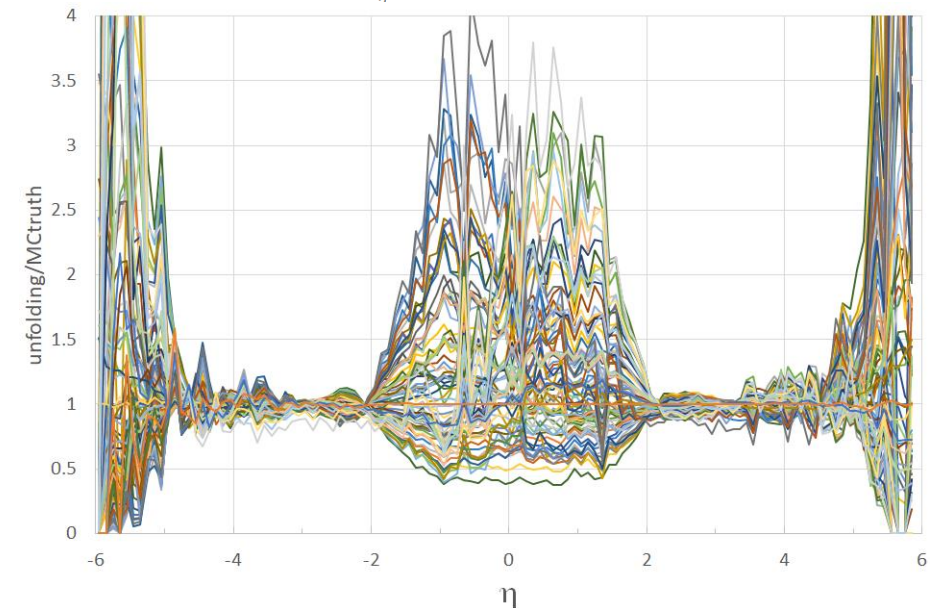
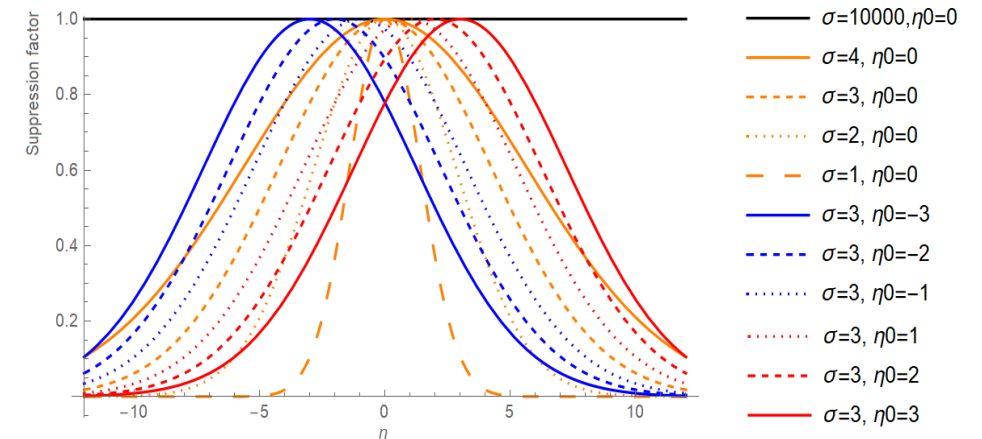


# Dependence on input distribution

- Distort simulated sample with suppression factor:

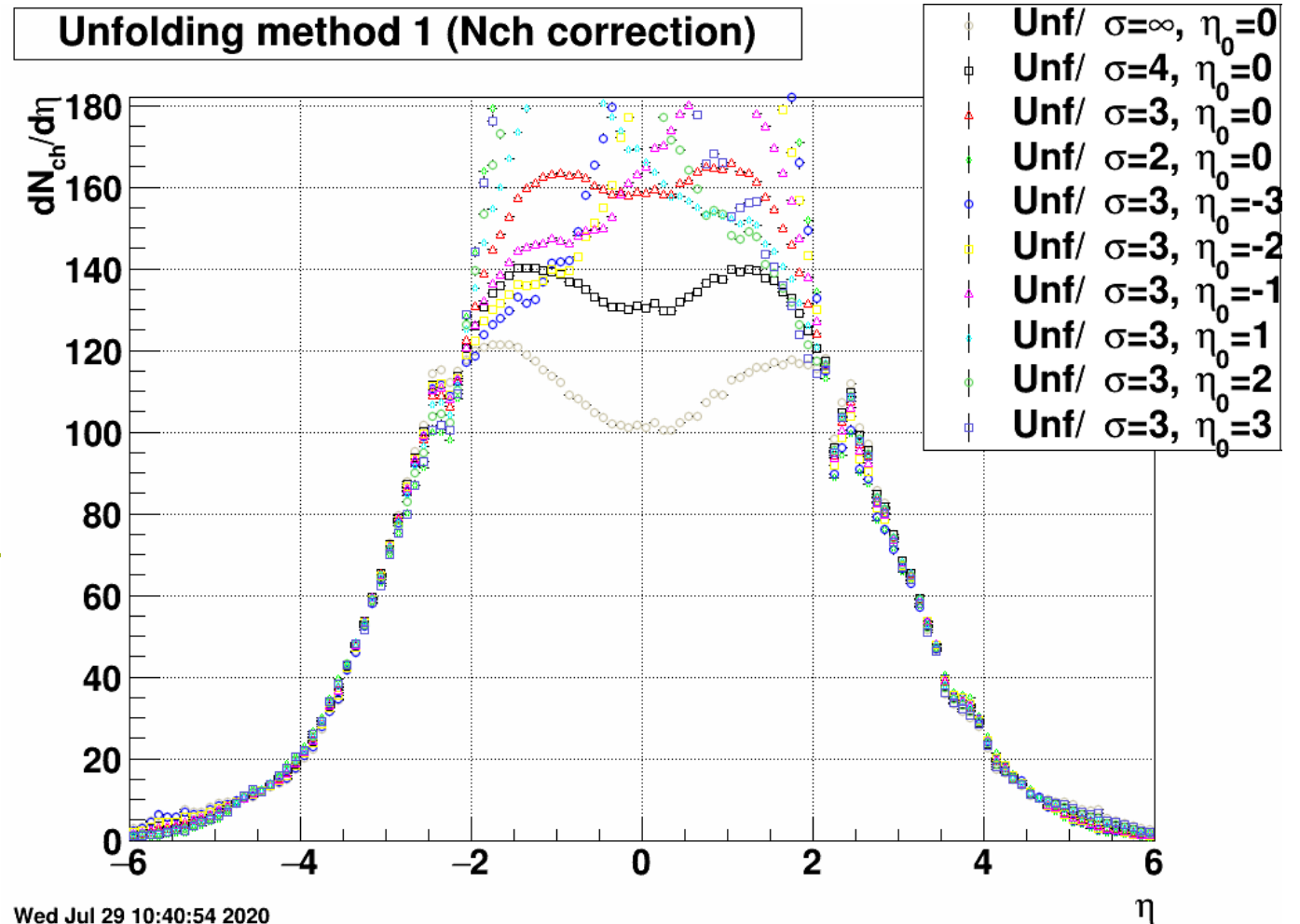
```
for(primary MC tracks with nonzero EPD hits)
{
  if(random_selection_based_on_Gaussian_distortion) continue;
  float eta = mctrack->Eta();
  MCtruth->Fill(eta);
  if(no_EPd_hits_for_this_primary) response->Miss(eta);
  for(EPD hits of this primary track)
  {
    int ring_bin = number between 0...31 (15&16 are the outermost rings of the two sides)
    response->Fill(ring_bin,eta);
  }
}
```

- Measure response with distorted sample
- Analyzed all combinations:
  - Unfold i-th sample with j-th distortion
  - If  $i=j$ : perfect unfolding
  - If distorting  $\sigma \approx 1$  or smaller: bad unfolding
  - Otherwise:  $\sim 10\%$  dependence in the EPD  $\eta$  region



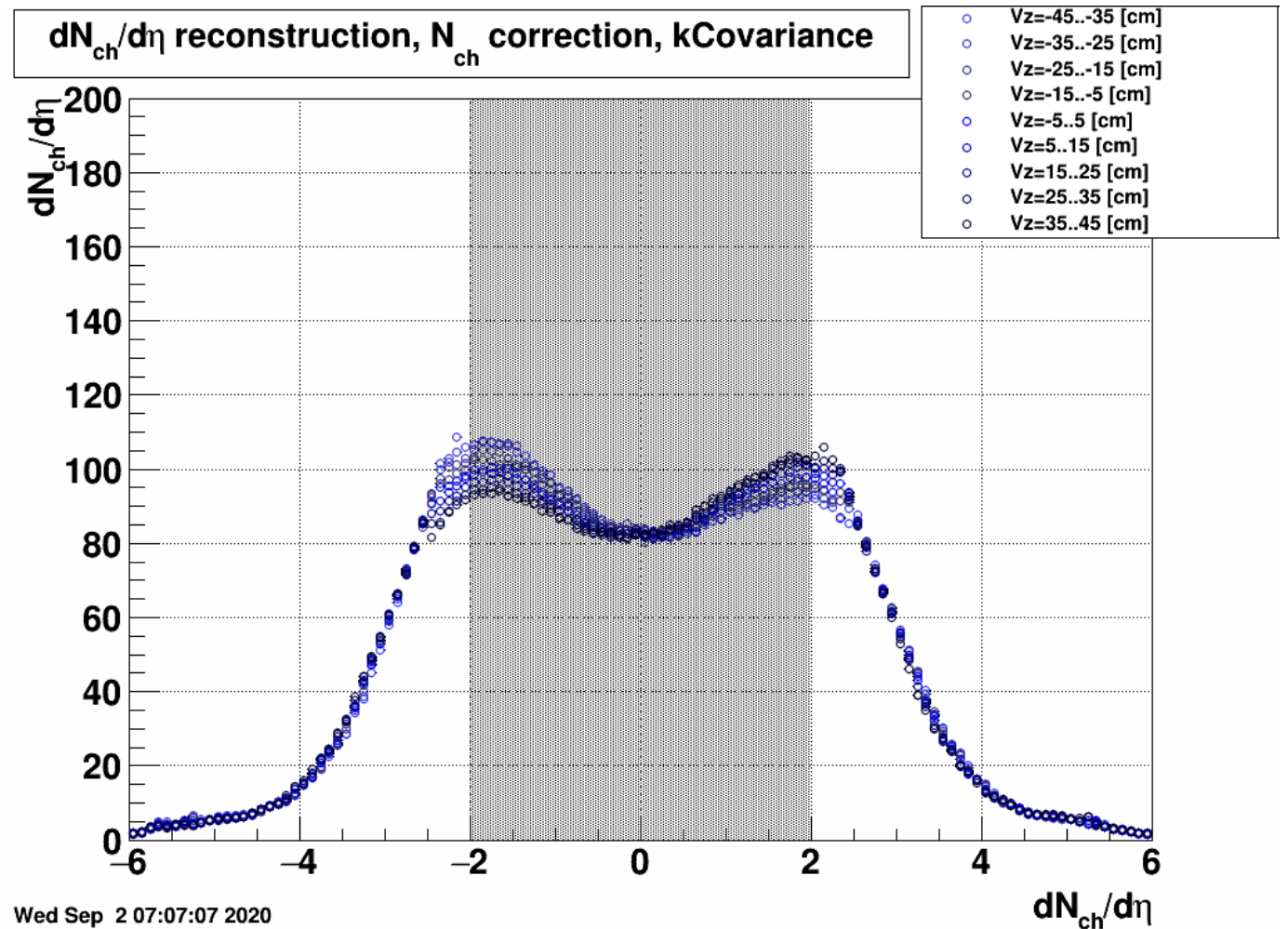
# Systematics: input $dN/d\eta$

- Most important systematic uncertainty: choice of input  $dN/d\eta$
- Huge uncertainty in the midrapidity region
- Mostly positive uncertainty: all distorted samples made distribution less wide
- Distortion that makes sample wider (i.e. rejecting midrapidity tracks) would yield negative uncertainty (i.e. lowering  $dN/d\eta$ )



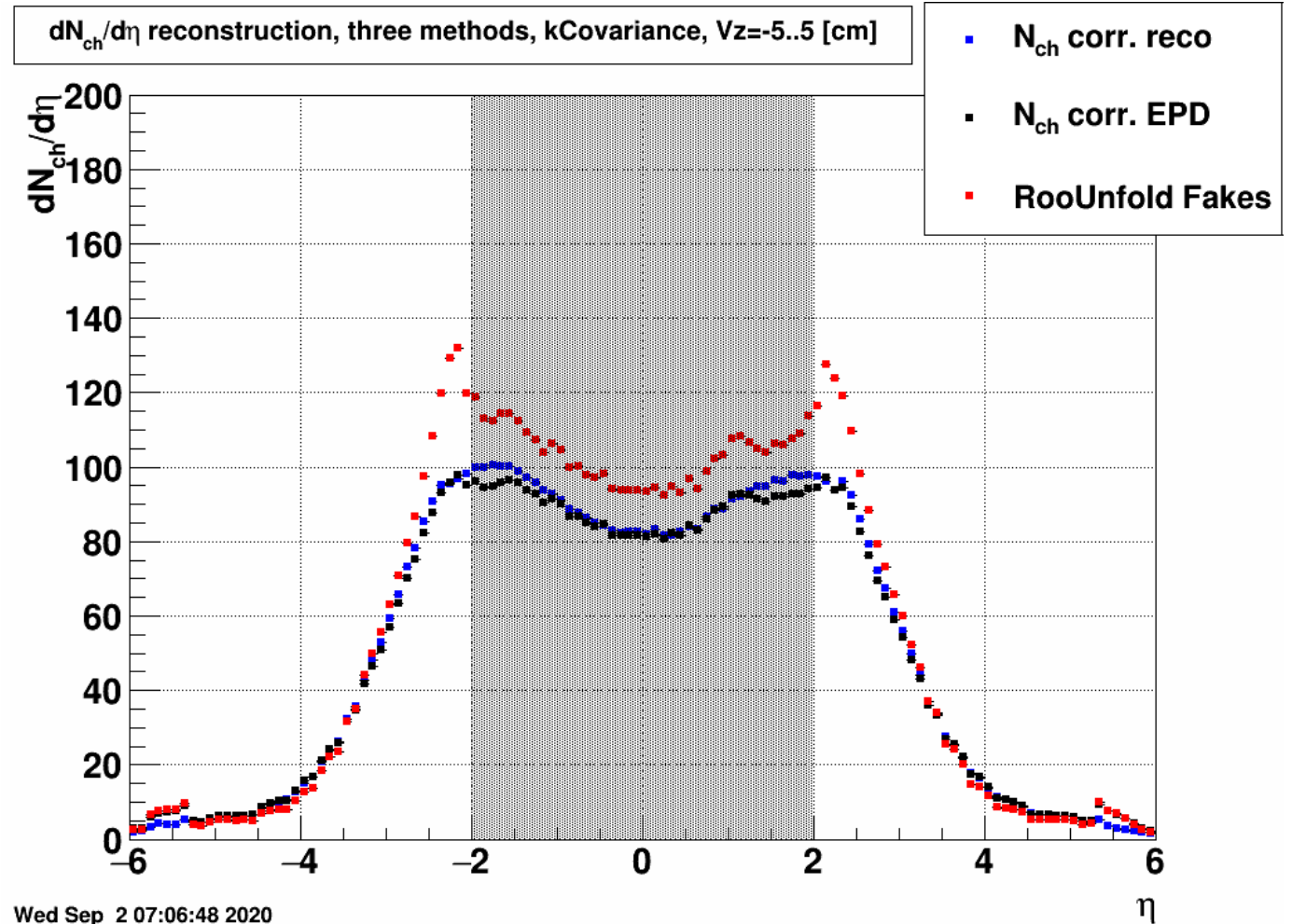
# Systematics: Vz binning

- Unfolded result depends on Vz bin
- Even if simulation also Vz-binned
- End result ( $dN/d\eta$ ) clearly needs to be Vz-independent
- Largest effect around  $|\eta| \approx 2$
- Differences to be included in systematics



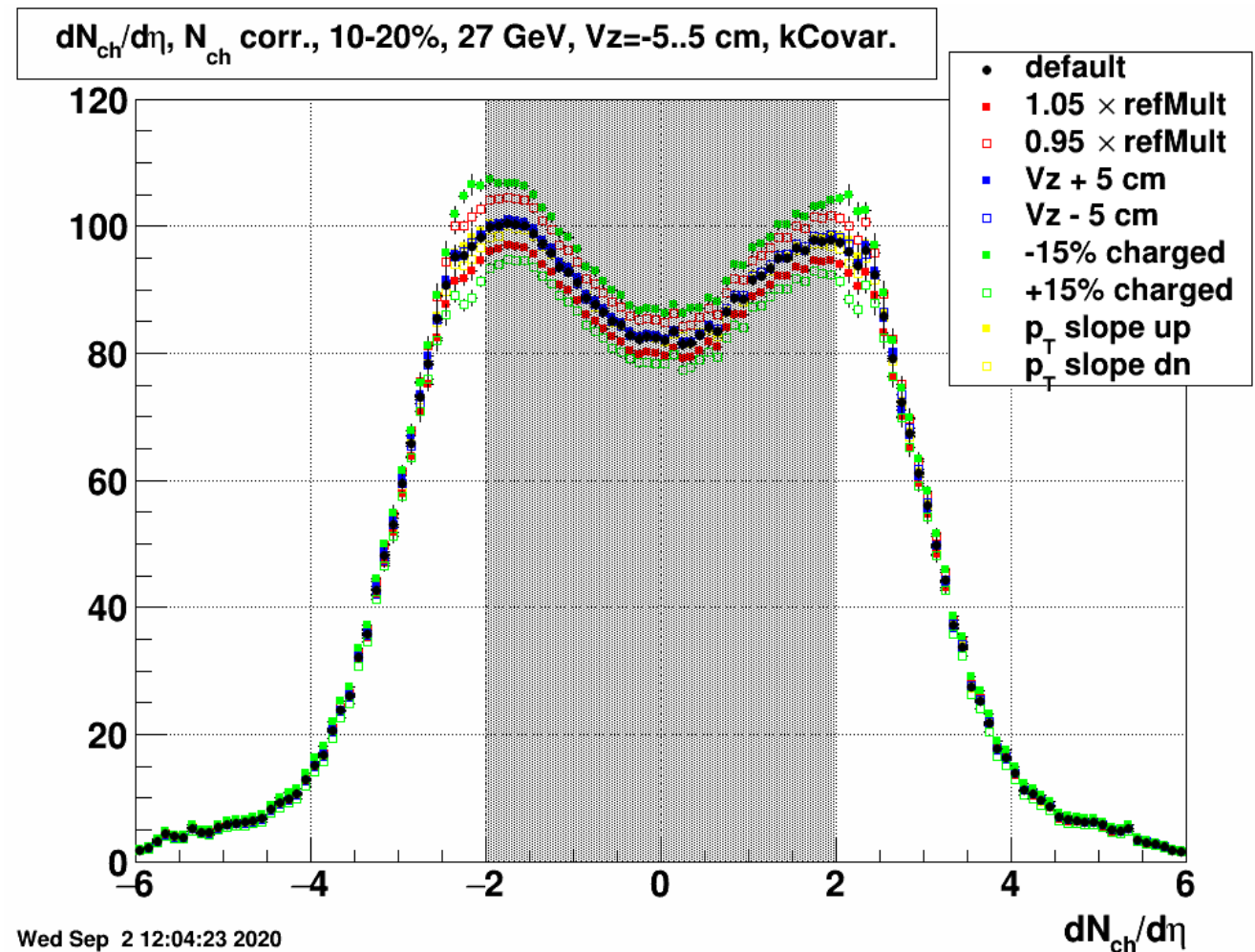
# Systematics: unfolding method

- Applied 3 different methods
  1. Unfolding  $dN/d\eta$ ; correcting via  $N_{\text{ch}}(\eta)/N_{\text{tot}}(\eta)$
  2. Correcting via  $N_{\text{ch}}(i_{\text{ring}})/N_{\text{tot}}(i_{\text{ring}})$ ; unfolding "corrected" EPD distribution
  3. Use RooUnfold's "Fakes" (neutrals  $\Leftrightarrow$  "fake" hits)
- "Fakes" different from the others
  - Also least reliable in terms of dependence on input  $dN/d\eta$
  - Reason of this unclear yet
- Other two methods match nicely



# Systematics: centrality, pT, charged ratio

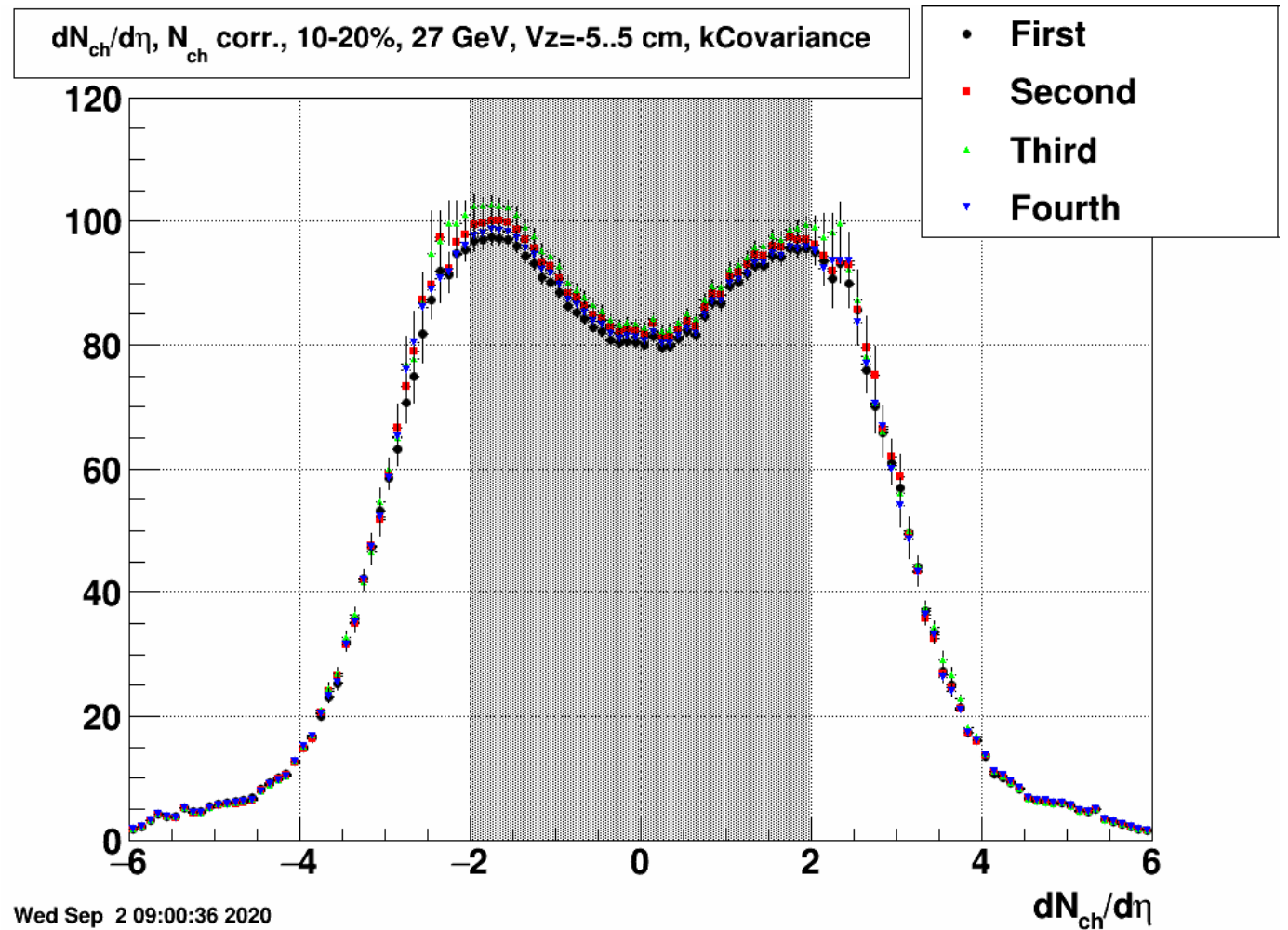
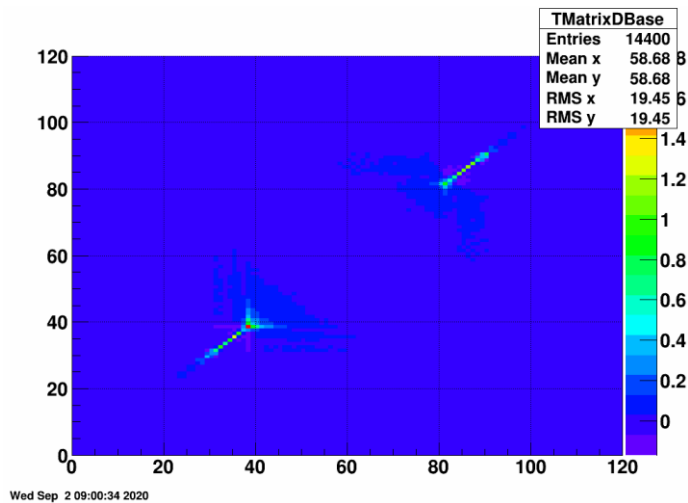
- Other sources of systematic uncertainty:
  - Centrality definition
  - Vz determination (this differs from Vz choice!)
  - Fraction of charged particles
  - pT slope of input sample
- These have a smaller effect than sources shown on previous slides
- Combined, still non-negligible





# Statistical uncertainties

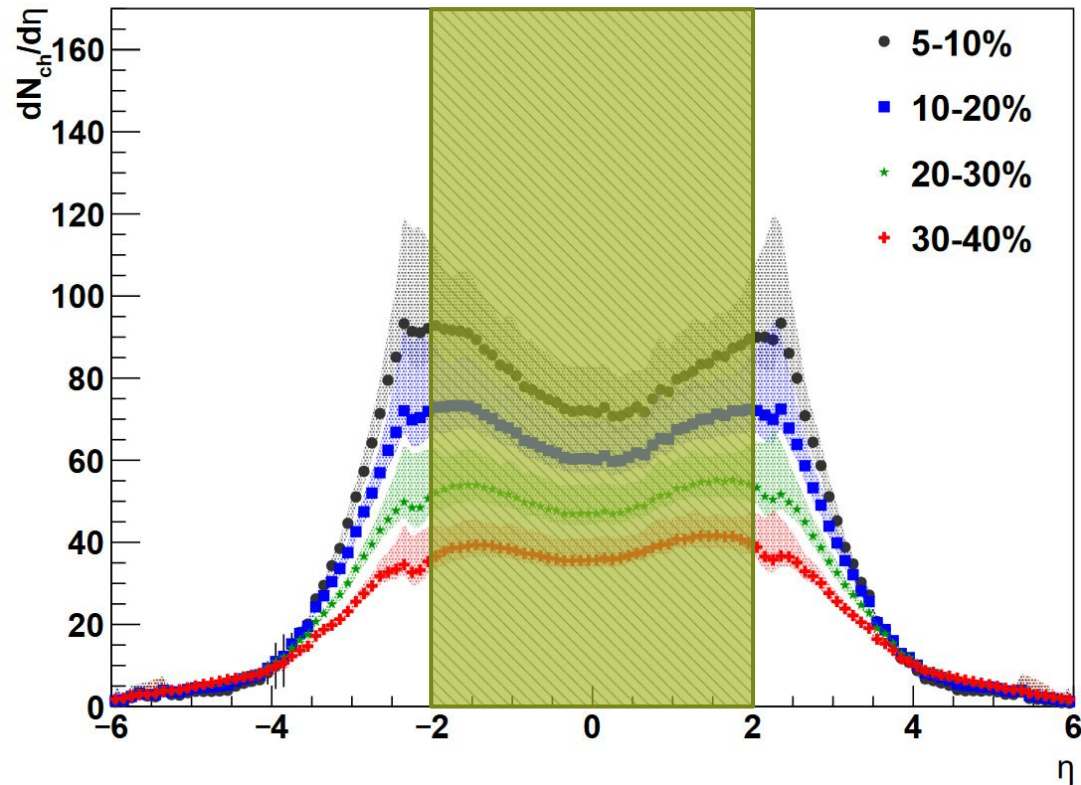
- Are statistical uncertainties of  $dN/d\eta$  datapoints reliable?
- Divided data sample into four sub-samples; these have a reasonable matching confidence level
- Covariance also available in ROOT:



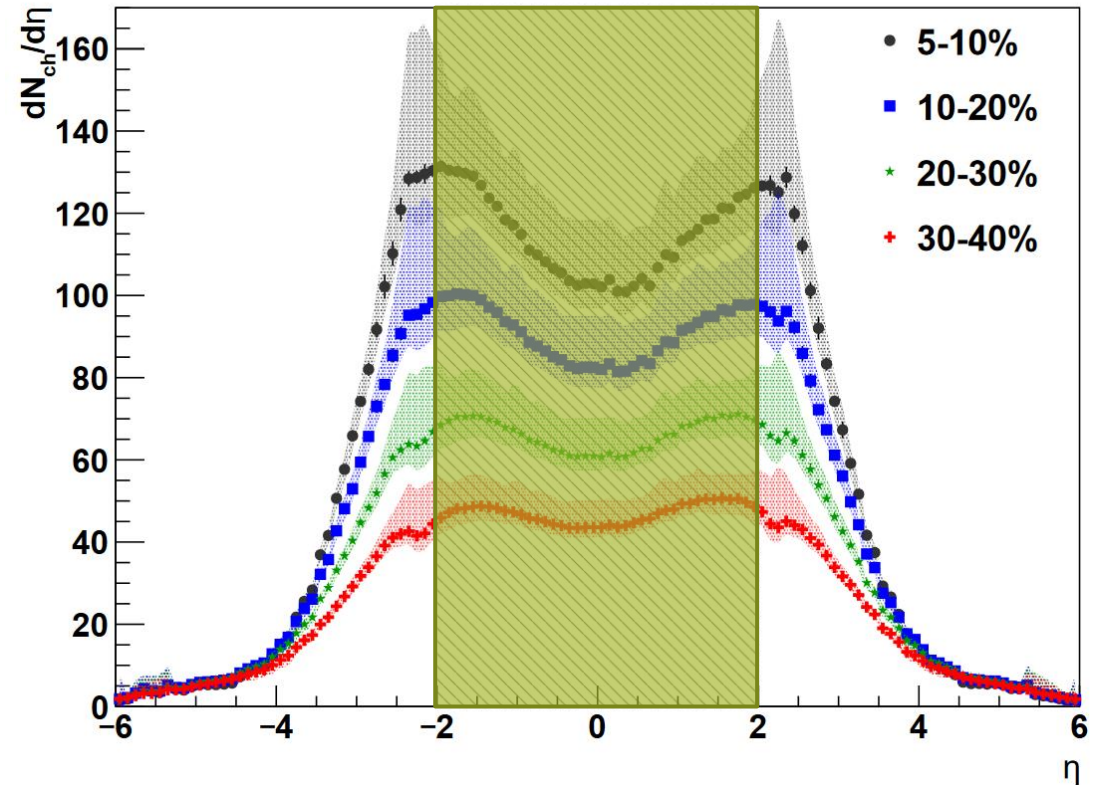
# Results at 19.6 & 27 GeV

- Unfolded results plotted with major systematic uncertainty sources
- Uncertainty from input  $dN/d\eta$ : huge for  $|\eta| < 2$ , region shaded out

STAR Au+Au @  $\sqrt{s_{NN}} = 19.6$  GeV,  $|v_z| < 5$  cm



STAR Au+Au @  $\sqrt{s_{NN}} = 27$  GeV,  $|v_z| < 5$  cm



# Summary

- Analysis mature, based on “provisional” data
- Systematic uncertainties considered:
  - $V_z$  determination
  - Centrality determination
  - $V_z$  choice (+40 & -40 cm compared)
  - Unfolding method
  - Charged/neutral ratio of training sample
  - $p_T$  slope of training sample
- Systematic uncertainty from input  $dN/d\eta$ 
  - In the  $|\eta| < 2$  region: huge uncertainty, region “shaded out” on plots
- Statistical uncertainties: very small, known covariance

